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**Formal GNS construction and WKB expansion in deformation quantization.** (English)

[Zbl 1166.53321](#)

Sternheimer, Daniel (ed.) et al., Deformation theory and symplectic geometry. Proceedings of the Ascona meeting, Switzerland, June 17–21, 1996. Dordrecht: Kluwer Academic Publishers (ISBN 0-7923-4525-8/hbk). Math. Phys. Stud. 20, 315-319 (1997).

From the text: The concept of deformation quantization has been defined and exemplified in *F. Bayen, M. Flato, C. Fronsdal, A. Lichnerowicz and D. Sternheimer* [Ann. Phys. 111, 61–110, 111–151 (1978; [Zbl 0377.53024](#), [Zbl 0377.53025](#))]. The existence of star-products on every symplectic manifold has been established in *M. De Wilde* and *P. B. A. Lecomte* [Lett. Math. Phys. 7, 487–496 (1983; [Zbl 0526.58023](#))] and in *B. V. Fedosov* [J. Differ. Geom 40, No. 2, 213–238 (1994; [Zbl 0812.53034](#))]. This gives a reasonable physical picture of the noncommutative algebra quantum observables with built-in classical limit. However, the discussion of a formal analogue of representations of the deformed algebra in some ‘Hilbert space’ seems to have been restricted to examples in the literature up to now. In [Commun. Math. Phys. 195, No. 3, 549–583 (1998; [Zbl 0989.53057](#))] the present authors have proposed how to construct formal pre-Hilbert spaces for the deformed algebra in the same category by means of a generalized Gel’fand-Naimark-Segal (GNS) construction. We now briefly review this construction and apply this in the next section to the Weyl star-product on the cotangent bundle of  $\mathbb{R}^n$  (which will give back the usual Schrödinger representation together with the Weyl symmetrization rule; details thereof can be found in the authors’ paper (loc. cit) .

The third section contains new material: we describe how to incorporate the usual WKB expansion into the framework of star-products and GNS representations by means of a certain positive linear functional on the deformed algebra having support on a projectable Lagrangian submanifold graph ( $dS$ ) of  $T^*\mathbb{R}^n$ . The main trick is to use a suitable form of the star-exponential  $e^{*\frac{i\hbar}{\lambda}S}$ .

For the entire collection see [[Zbl 0923.00023](#)].

**MSC:**

- [53D55](#) Deformation quantization, star products
- [81Q20](#) Semiclassical techniques, including WKB and Maslov methods applied to problems in quantum theory
- [81S10](#) Geometry and quantization, symplectic methods

Cited in **6** Documents

**Full Text:** [arXiv](#)